

Hi-C Observations of Penumbral Bright Dots: Comparison with the IRIS Results

Shane Alpert ^{1,2}

(S. K. Tiwari ², R. L. Moore ², S. L. Savage ², A. R. Winebarger ²)

1) Dept. of Physics and Astronomy, Rice University

2) NASA Marshall Space Flight Center



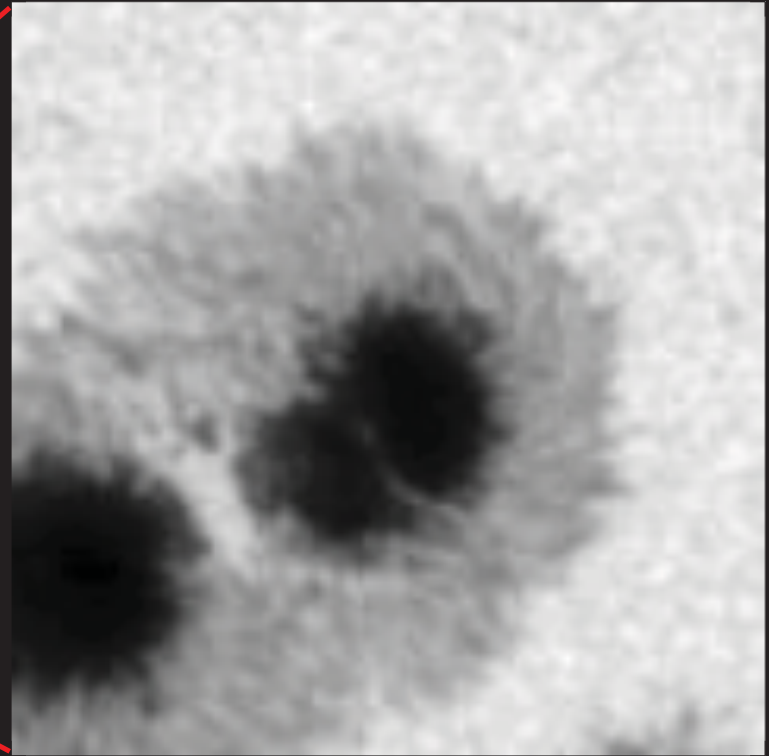
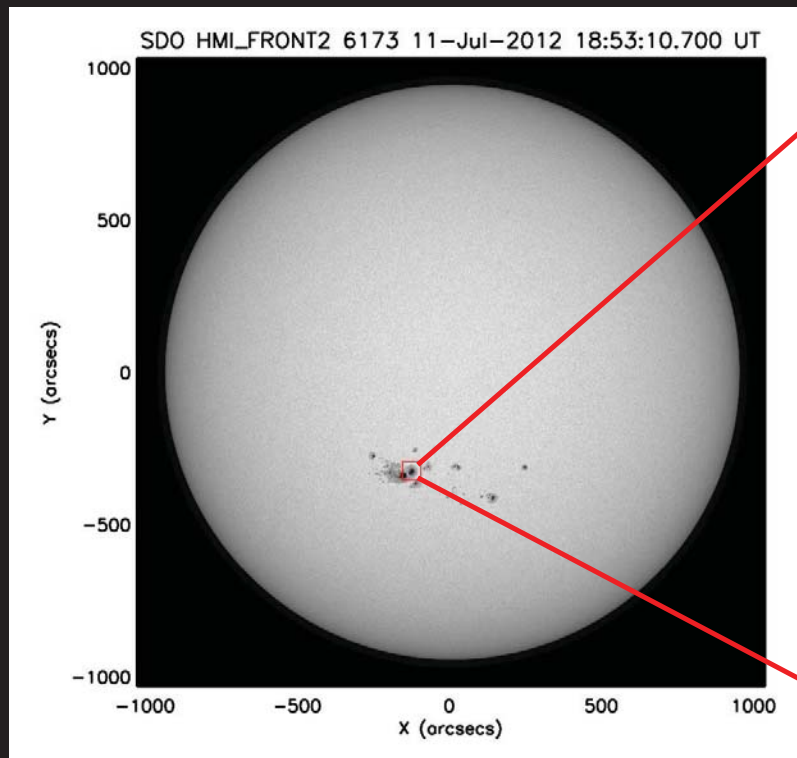
LWS Workshop – Nov. 3, 2014



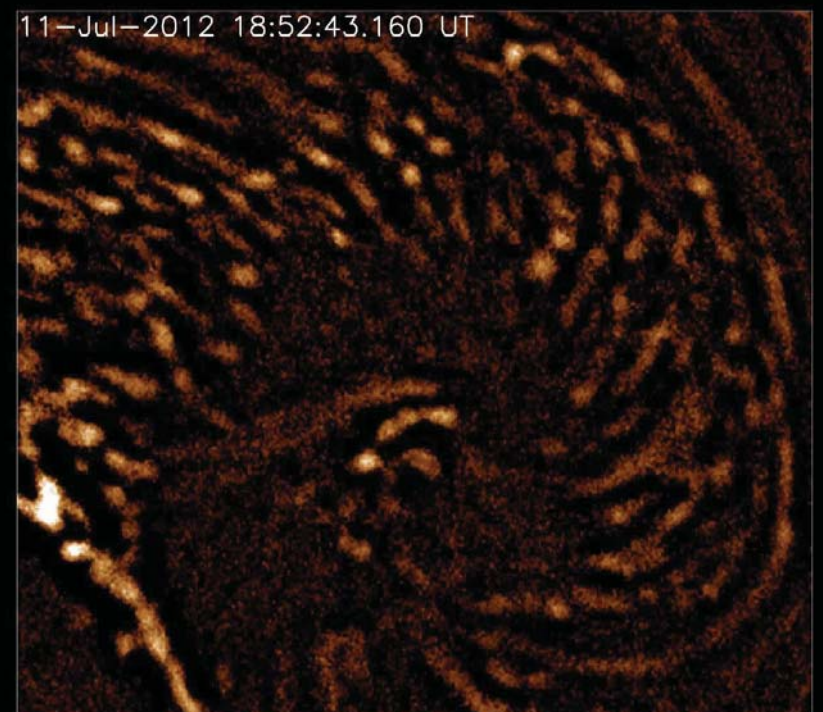
Overview

- Penumbral bright dots in brief
- Qualitative description
- Quantitative properties
- Comparison to IRIS observations
- Possible causes

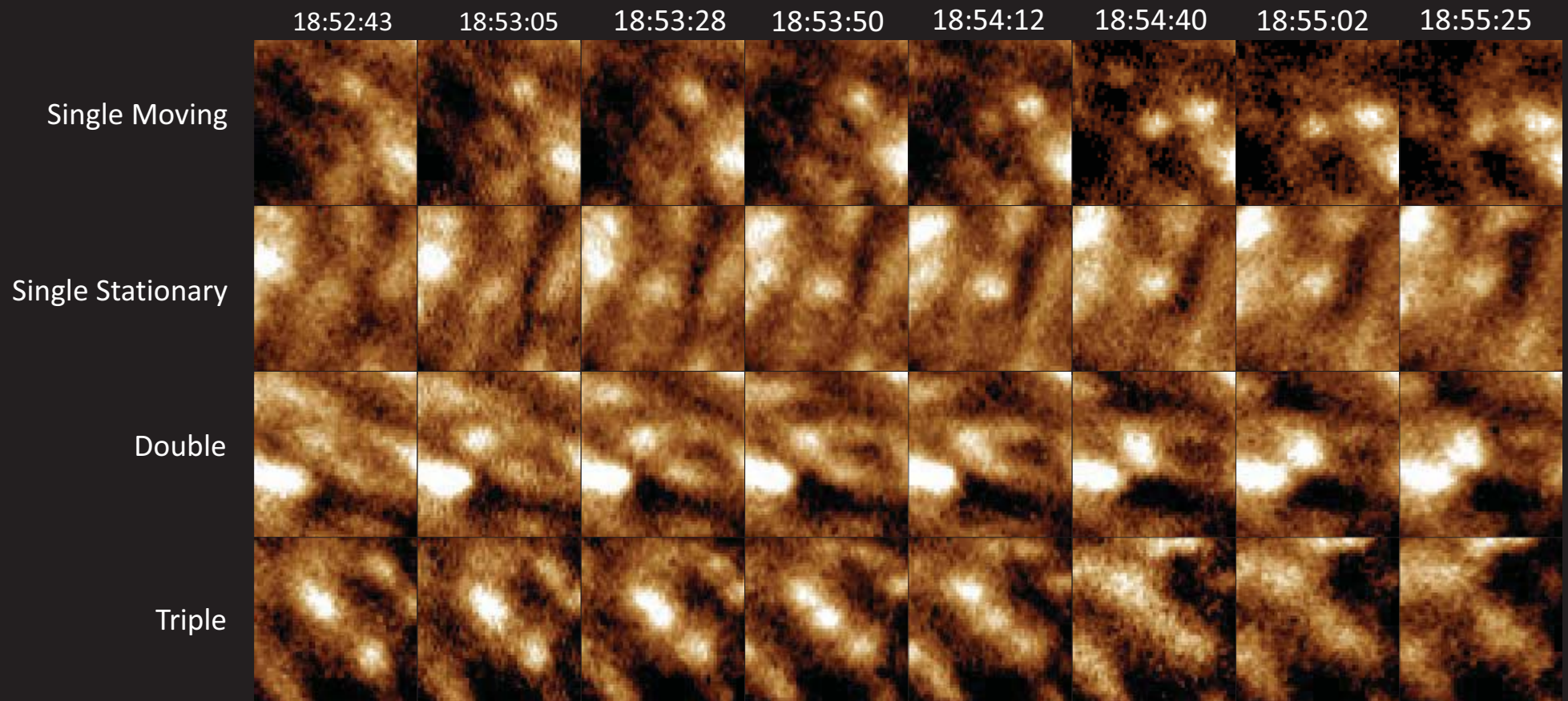
Penumbral Dots in Brief



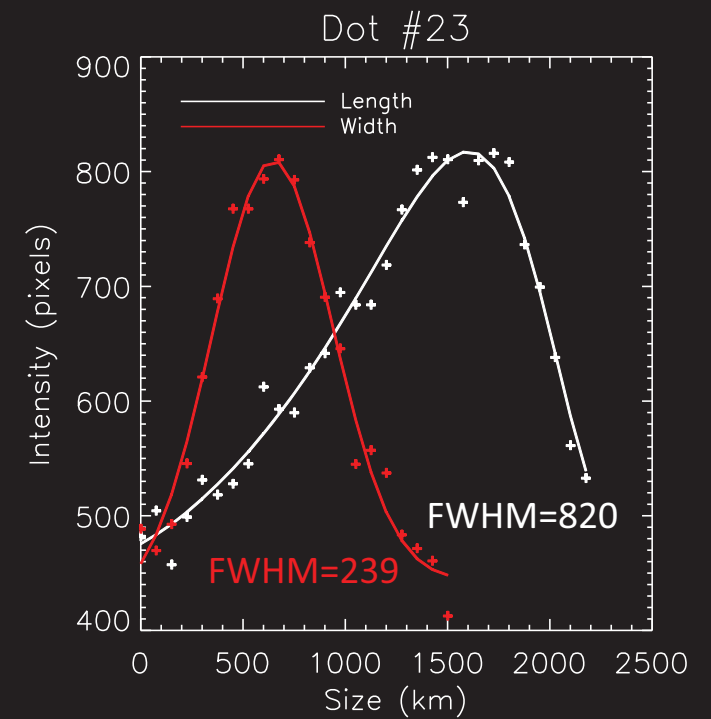
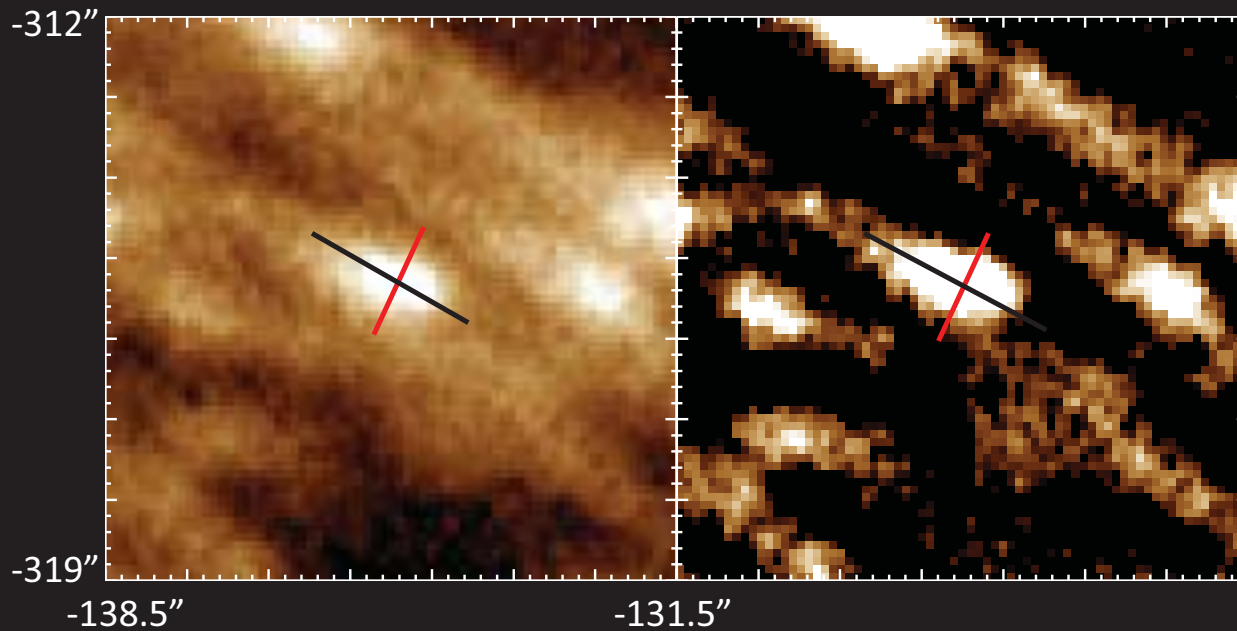
Qualitative Description



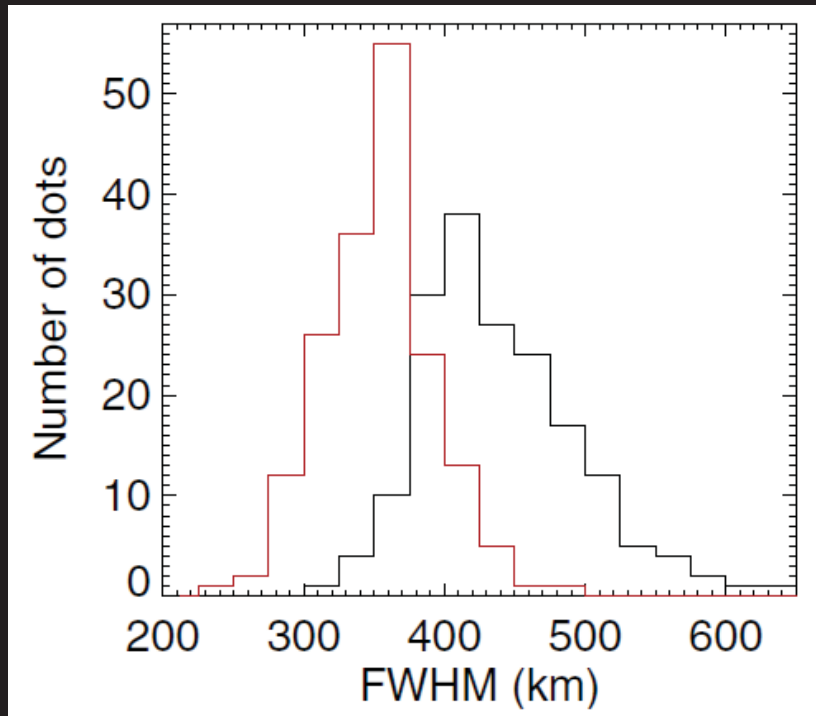
Qualitative Description



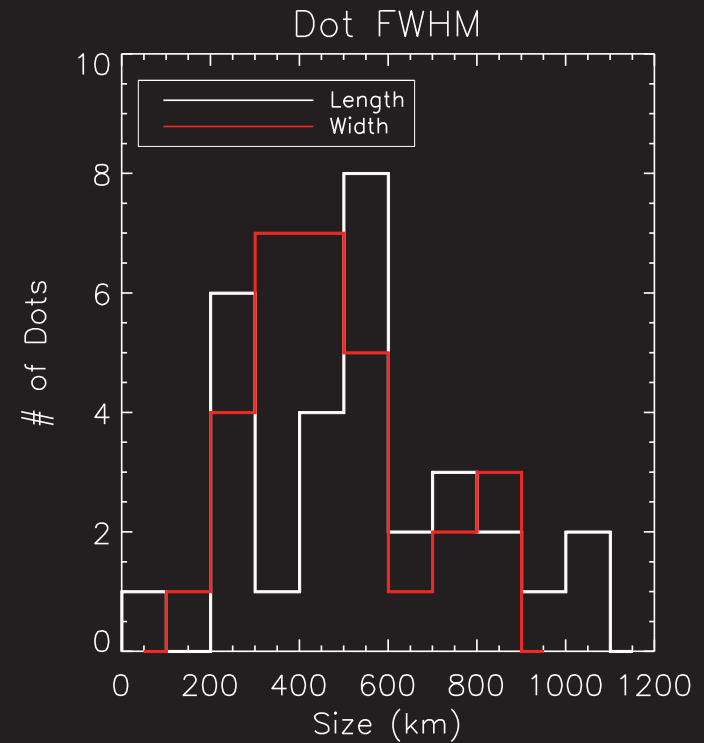
Quantitative Properties - Size



Quantitative Properties - Size

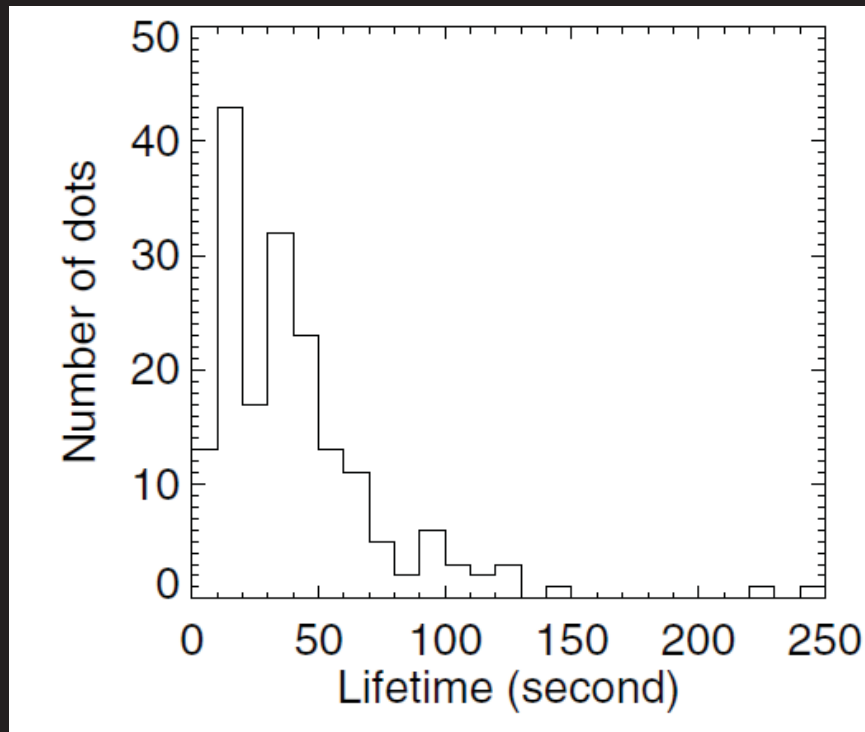


Tian et al. (2014) ApJ, 790, L29

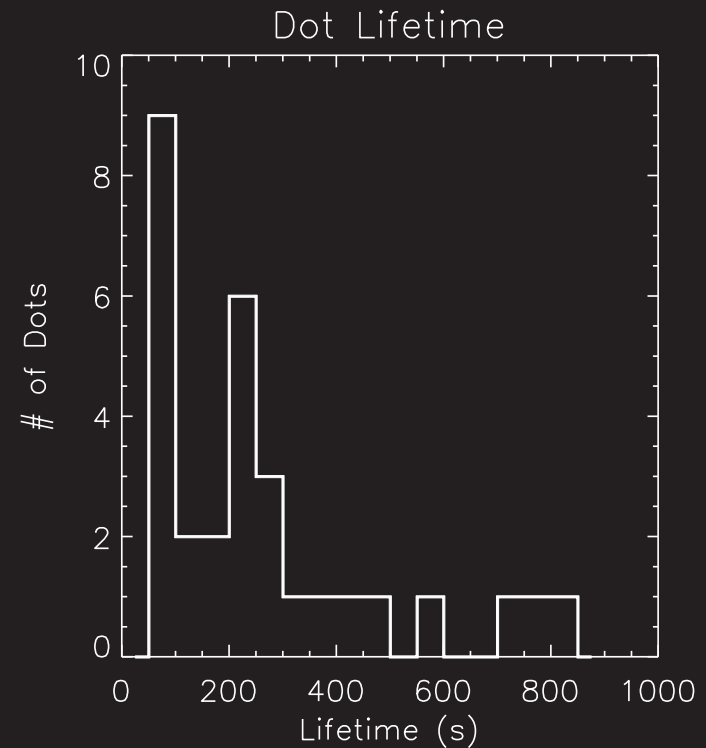


Alpert et al. (2014) (pending)

Quantitative Properties – Lifetime

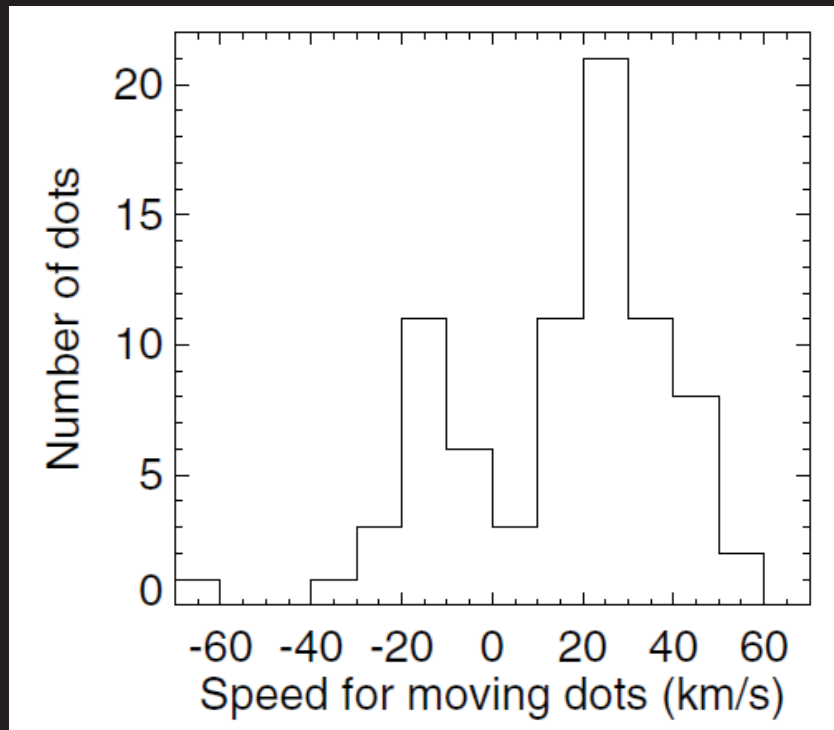


Tian et al. (2014) ApJ, 790, L29

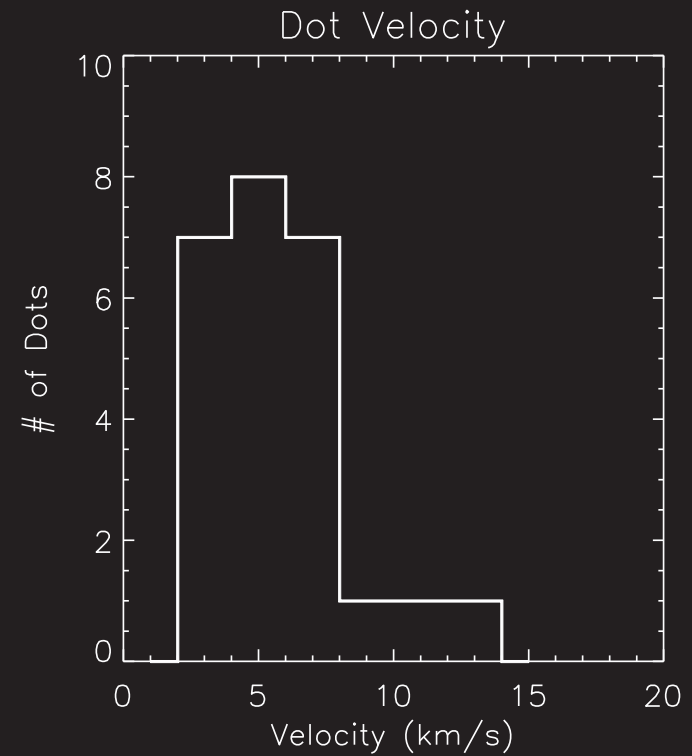


Alpert et al. (2014) (pending)

Quantitative – Velocity

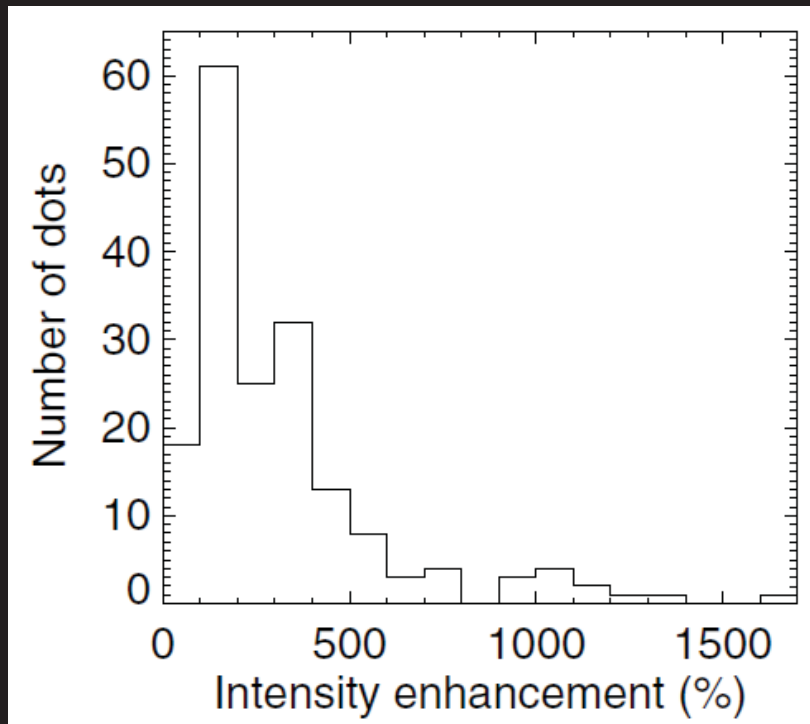


Tian et al., 2014, ApJ, 790, L29

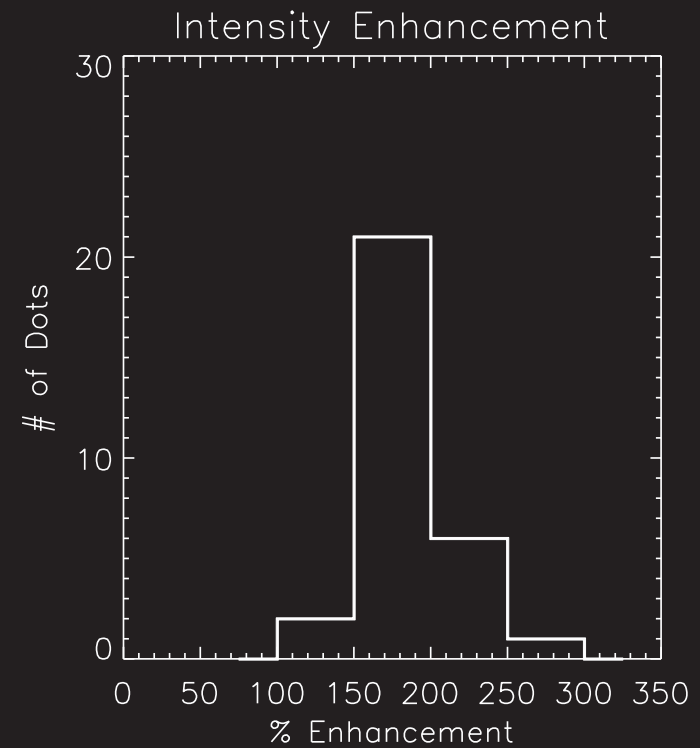


Alpert et al. (2014) (pending)

Quantitative Properties – Intensity

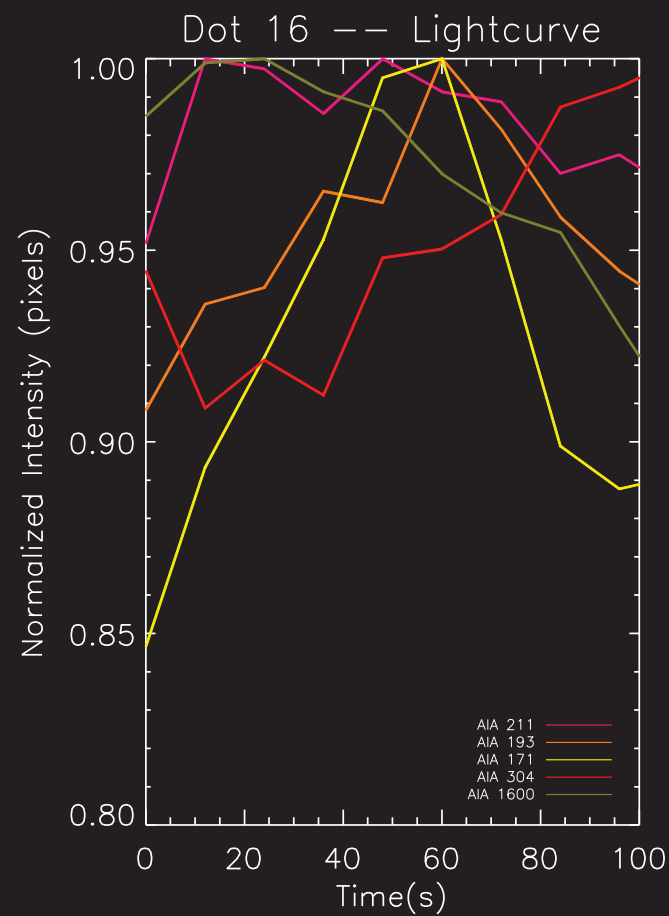
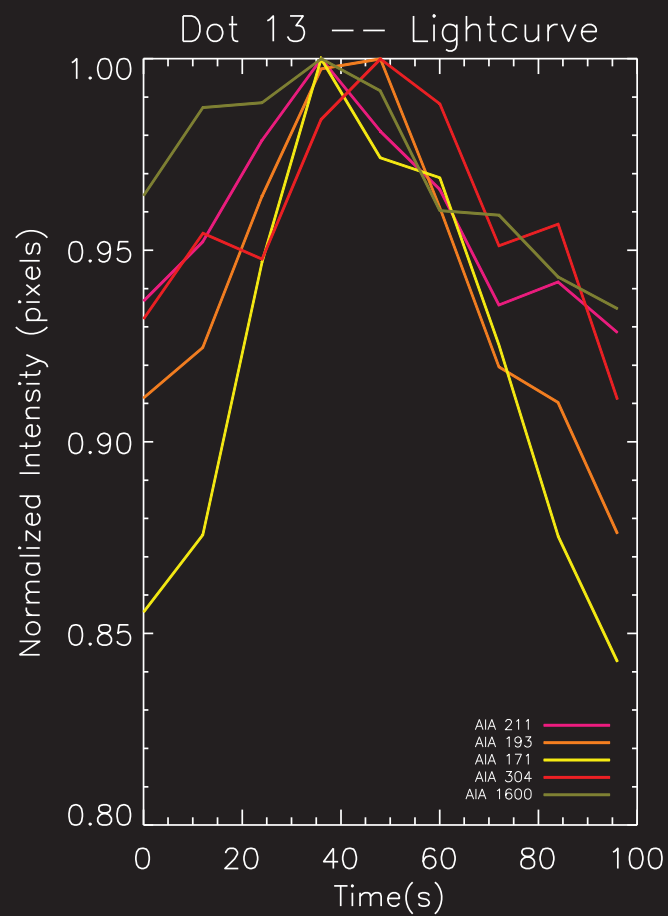


Tian et al. (2014) ApJ, 790, L29



Alpert et al. (2014) (pending)

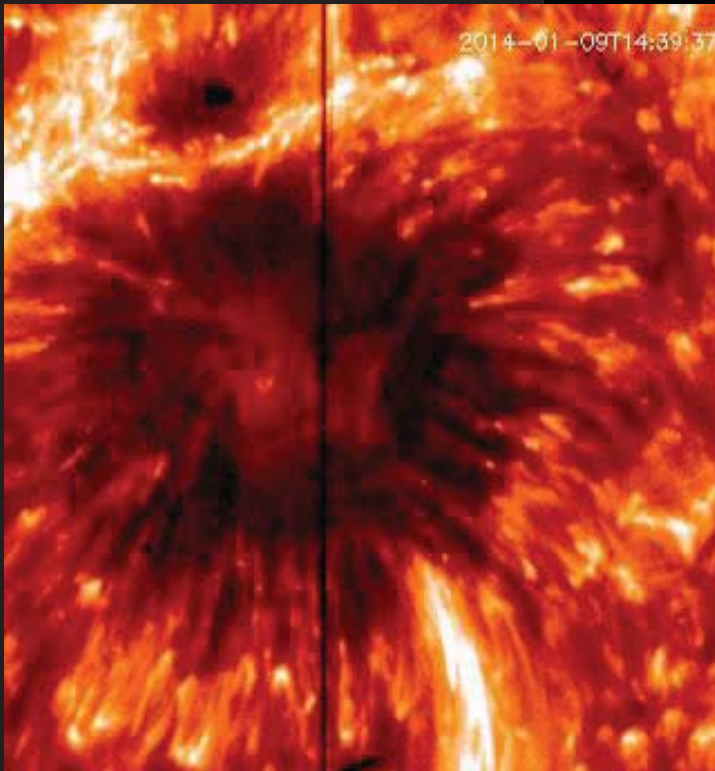
Quantitative – Lightcurves



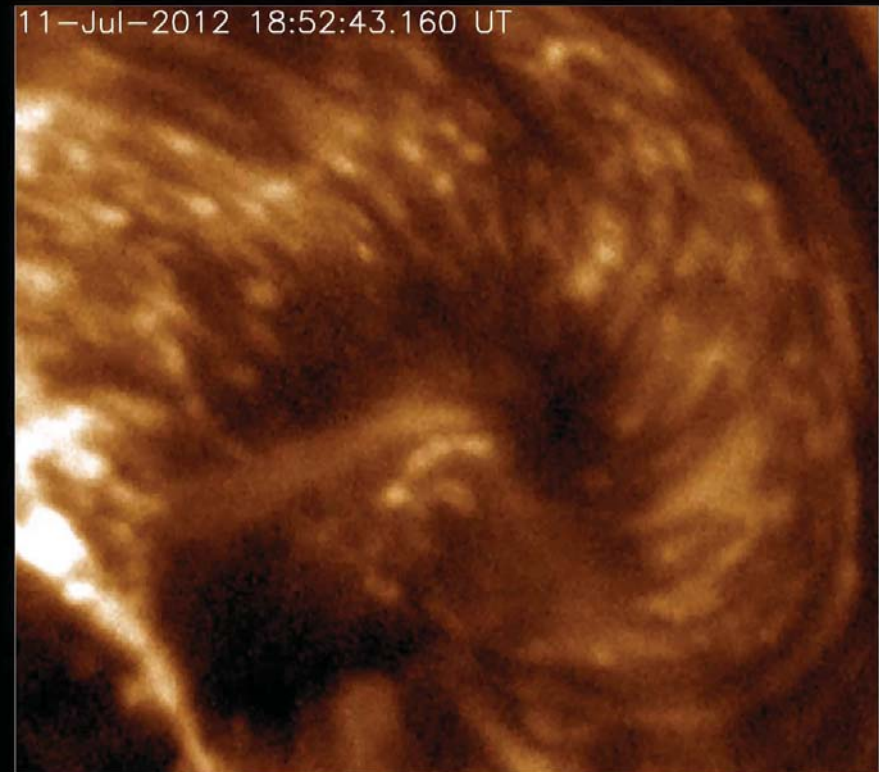
Comparison to IRIS Observations

	Tian et al. (IRIS)	This research (Hi-C)
Length (km)	439.0	557.7
Width (km)	352.3	472.4
Lifetime (s)	41.3	268
Speed (km/s)	N/A	5.8
Intensity Enhancement (%)	319.4	188

Comparison to IRIS Observations



Tian et al., 2014, ApJ, 790, L29



Possible Causes

- Plasma downflow/falling plasma
- Repeated reconnection → nanoflares
- Upflows
- Oscillatory motion due to waves

Future Work

- Examine lightcurves
- Larger sample of dots
- Numerical modeling to test possible mechanisms

Thank you!

